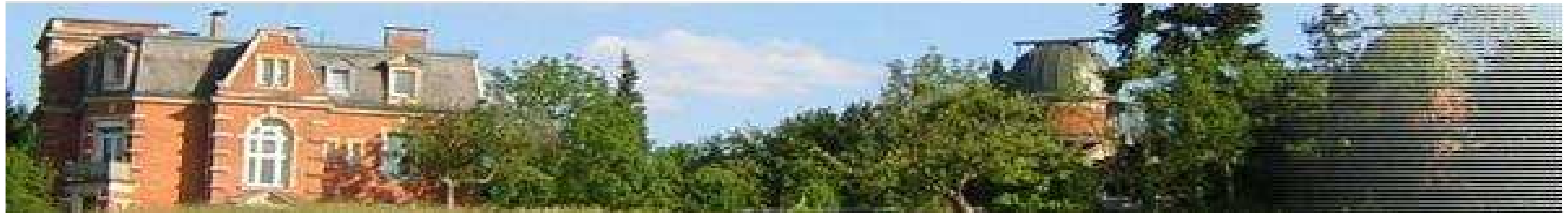


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DETAIL/SURFACE

Exercises

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Non-LTE line-formation calculations with DETAIL & SURFACE

I) Directory Structure (in **group**)

DETSURF

atmospheres	atmospheric structures in DETAIL format
atomic_data	atomic input data for Exercise 2
exercises	exercise material (pdf)
input	DETAIL/SURFACE control files
runs	working directory



II) Exercise 1: Invoking DETAIL & SURFACE

Task: run DETAIL & SURFACE

enter **runs**

sample control files are

DETAIL_input_bss_vega
SURFACE_input_bss_vega

log O/H+12=8.53

data required by DETAIL to solve SE and RT for

- O I model atom similar to one by Baschek, Scholz & Sedlmayr, A&A, 55, 375 (1977) **BSS**
- model atmosphere for Vega adopted from Kurucz' homepage

and by SURFACE to calculate emergent flux using

- NLTE level populations
- refined line-broadening



detail < control_file > output

- requires possibly auxiliary files, e.g. photoionization cross-sections
- **output**: detailed information, required for e.g. model atom verification or physical interpretation
→ previous discussion
- NLTE level populations: **fort.7**

surface < control_file > output

- requires possibly auxiliary files, e.g. broadening tables
- **output**: detailed information, as above
→ previous discussion
- physical flux: **fort.7**
- rectified flux: **fort.70**



Comments

- last call of **ABUNDANCE** sets values for calculation
- abundances are normalised to H+He !!!
- **BABS**: $b_i = n_i/n_i^{\text{LTE}}$ (Zwaan) else $b_i = \frac{n_i/n_i^{\text{LTE}}}{n_c/n_c^{\text{LTE}}}$ (Menzel)
- convergence criterion: $\Delta n/n \sim 10^{-4}$
- inner boundary condition: LTE ($b_i=1$) $\longrightarrow \tau_{\text{Ross}} \sim 100\text{-}200$ sufficient
- error handling by DETAIL/SURFACE - but: user mistakes possible
- user has to provide the correct **ODF file**
via `ln -sf ~/DETSURF/ODF/xyz ODF`
- meaningful renaming of output files recommended

verify your output by comparing with **~/DETSURF/solutions: vega*bss***



III) Exercise 2: O I model atom

$\log O/H+12=8.53$

Task:

construction of a simple model atom for O I,
similar to that of Carlsson & Judge, ApJ, 402, 344 (1993) **CJ**

- energy levels up to $n=4$ shall be considered, i.e. also excited singlet and triplet states, with more up-to-date transition data
- all required data to be found in **atomic_data** preformatted for use with DETAIL & SURFACE
- use **DETAIL_input_bss_vega** and **SURFACE_input_bss_vega** (in **runs**) and the DETAIL/SURFACE manual for guidance
- calculate NLTE populations with DETAIL (check convergence!) & emergent flux with SURFACE for the Vega model

verify your results by comparing with **~/DETSURF/solutions/vega*cj***,
sample input files in **~/DETSURF/solutions/input: DETAIL_input_cj_vega,**
SURFACE_input_cj_vega



IV) Exercise 3: NLTE line-formation with DETAIL & SURFACE

- Task:
- a) NLTE calculations for 3 model atoms for O I for Vega and a Vega-like supergiant
 - b) comparison with LTE
 - c) determination of NLTE abundance 'corrections'
 - d) NLTE & LTE profiles @ same W_λ
 - e) physical interpretation

- 3rd O I model atom to be used: **PBBKV**
Przybilla, Butler, Becker, Kudritzki & Venn, A&A, 359, 1085 (2000)

control files are in **input: *pbbkv***

there, also additional control files for the SG can be found
- copy these to **runs**



Grotrian diagrams

model atoms

PBBKV

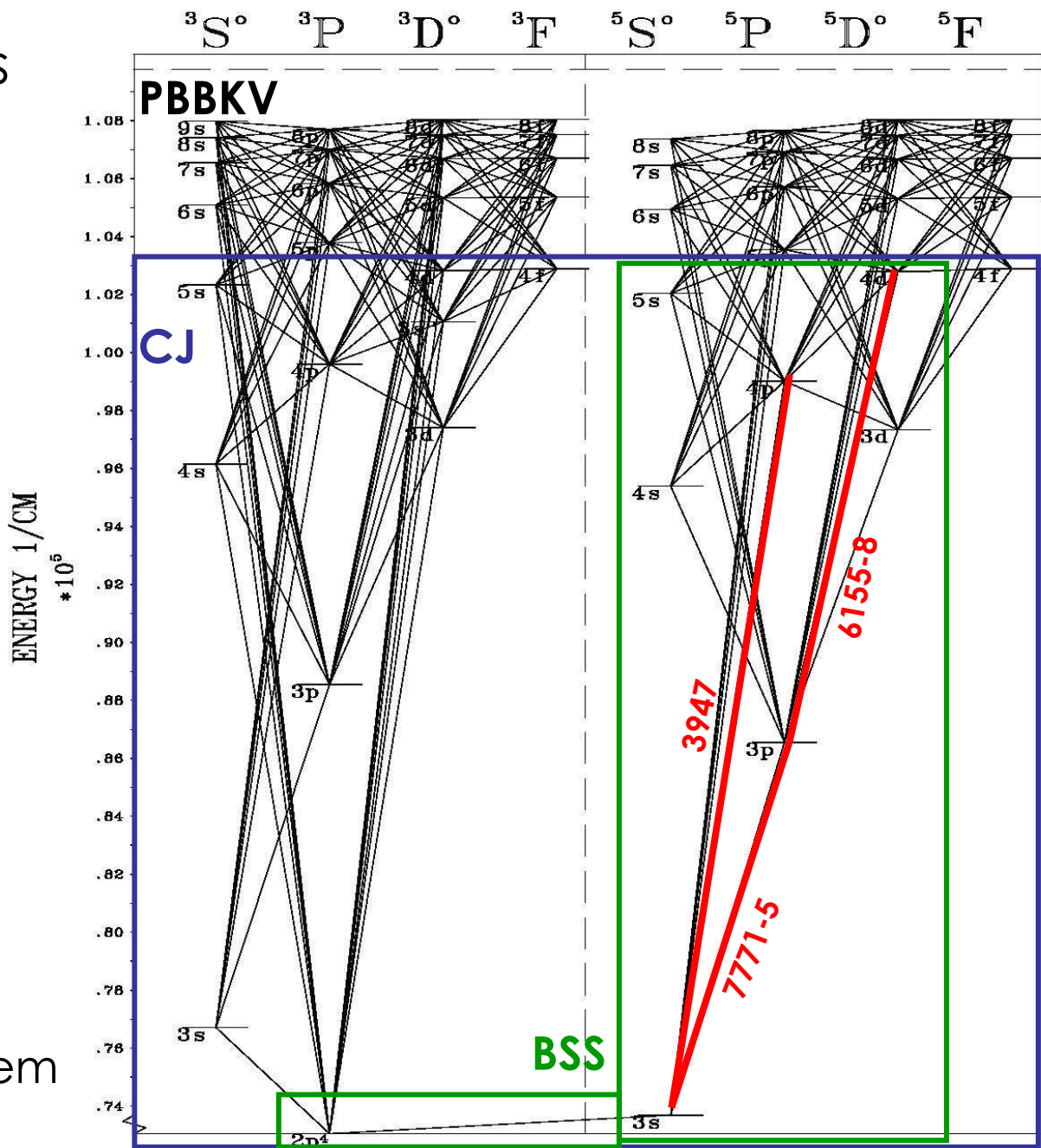
CJ

BSS



increasing complexity

+ singlet spin system



3a) NLTE calculations for 3 O I model atoms

- run all control files to generate 6 emergent NLTE fluxes
check for convergence – use 120 iterations for **CJ** & **PBBKV**
 - compare NLTE line profiles
 - I $\lambda\lambda 3947$
 - I $\lambda\lambda 6155-8$
 - I $\lambda\lambda 7771-5$
- from **BSS**, **CJ** & **PBBKV** model atoms for Vega & SG
use `n1tep1ots_f1`
- your interpretation?



Intermezzo: Visualisation with the Fawlty Language

If you are familiar with IDL, you will feel at home. Note that FL is a project under development, so don't expect any functionality beyond the most basic. There is no manual for FL at present.

invoke the FL by: `f1` exit via: `exit`

- reading of SURFACE **fort.70** (**fort.7**) files:

```
read_surf_f1, 'filename', lambda, flux
```

- plotting of flux vs. lambda:

```
plot, lambda, flux, xr=[7771,7776], yr=[0.1,1.05], xsty=1, ysty=1
```

useful keywords: `xtitle=' '`, `yttitle=' '`, `charsize=1.0`, `lin=1`

- overplotting:

```
oplot, lambda2, flux2
```

- text output in plot:

```
xyouts, xcoordinate, ycoordinate, ' '
```



- reading atmospheric structure (in DETAIL format):

`read_st_fl, 'filename', m, t, na, ne`

m: mass scale
t: temperature
na: number density of heavy particles
ne: number density of electrons

- logarithmic plotting:

`plot, alog10(m), t`

- reading NLTE populations & departure coefficients (DETAIL **output** file)

`read_pops_fl, 'filename', n, b, lab`

n: NLTE level populations
b: departure coefficients
lab: labels of levels

- plotting departure coefficients in 3e) by:

`departures_fl`

edit filenames in **runs/departures_fl.pro**

generates PS file



3b) Comparison with LTE

- copy SURFACE control files **input/*LTE*** to **runs**
 - run all control files to generate 6 emergent LTE fluxes for same abundance $\log O/H+12=8.53$
 - compare LTE with NLTE line profiles for
 - I $\lambda\lambda 3947$
 - I $\lambda\lambda 6155-8$
 - I $\lambda\lambda 7771-5$
- from **PBBKV** model atom for Vega & SG
use `1tep1ots_f1`
- your interpretation?



3c) Determination of NLTE abundance 'corrections'

- due to time constraints no practical work here
- procedure:
 - construct LTE curve-of-growth $W_\lambda = f(O \text{ abundance})$
 - find abundance at which W_λ^{NLTE} is reproduced
 - difference of NLTE and LTE abundance
→ **NLTE abundance 'correction'**
- schematically for 2 lines in Vega & SG



3d) NLTE & LTE profiles @ same W_λ

- calculate LTE emergent fluxes for

log O/H=8.56 & 9.11 (Vega)

log O/H=8.72 & 10.10 (SG)

→ reproduce W_λ^{NLTE} (log O/H=8.53) for O I $\lambda 6158$ and O I $\lambda 7771$

- compare LTE line profiles for O I $\lambda 6158$ and O I $\lambda 7771$ with NLTE line profiles for log O/H=8.53 (**PBBKV** model) in Vega and SG
- your interpretation?



3e) Physical interpretation: a first step

- plot departure coefficients b_i for strategic levels
 - I ground state
 - I $3s\ ^5S^{\circ}, 3p\ ^5P, 4d\ ^5D^{\circ}$
 - II ground state
- as a function of $\log m$ for Vega & SG
use `departures_f1`
- your interpretation?

